

# DUNE CD1 Refresh: GLOBES Configurations, and Sensitivity studies

Daniel Cherdack  
Colorado State University

Combined LB Task Force Meeting  
Monday April 6<sup>th</sup>, 2015

# Outline

- GLoBES
- Study parameters
  - Assumptions
  - Oscillation parameters
  - Non-oscillation systematics
- GLoBES inputs
  - Fluxes
  - Cross sections
  - Smearing
  - Efficiencies
- Spectra comparisons
- Using this GLoBES setup

# Analysis Basics

- What do we measure?
  - Number of  $\nu$  interactions as a function of reconstructed neutrino energy
  - Convolution of (flux) x (osc. prob.) x (xsec) x (nuclear effects) x (efficiency) x (energy smearing)
- Hypothesis testing
  - Compare data to simulation (e.g.  $\chi^2$  test)
  - Adjust parameters of interest (oscillation parameters) until the simulation best describes the data (min  $\Delta\chi^2$ )
  - Also allow nuisance parameters to fluctuate
  - Determine likelihood that data is consistent/inconsistent with the null hypothesis
  - Determine probabilist confidence intervals for fit parameter values
- What to do in the absence of data?
  - “Make up” data based on simulations and specific values for parameter(s) of interest
  - Toy MC throws: Randomly fluctuate all free parameters and statistics within priors many, many times
  - Asimov data set: use the simulated prediction, i.e. the median simulation

# GLOBES

- These studies have a lot of free parameters
- There are a lot of measurements and scenarios to explore
- Fitting requires quickly propagating parameter changes to spectral differences
- How do we quickly generate spectra for each scenario, and set of free parameters?
  - Normally build simulations event-by-event
  - GLOBES produces binned spectra based on inputs by convolving all the inputs (nuclear effects included in xsec or energy smearing)
  - Allows the production of spectra without detailed simulations
  - Very fast, but at the expense of some detail
- We have detailed simulations to construct GLOBES inputs

# Study Parameters: Basics

- Beams:

- Reference beam
  - Current “best” beam
  - Expected optimization
  - Perfect focusing
  - Power and POT (needed for GLoBES)
- } See Beam Talk

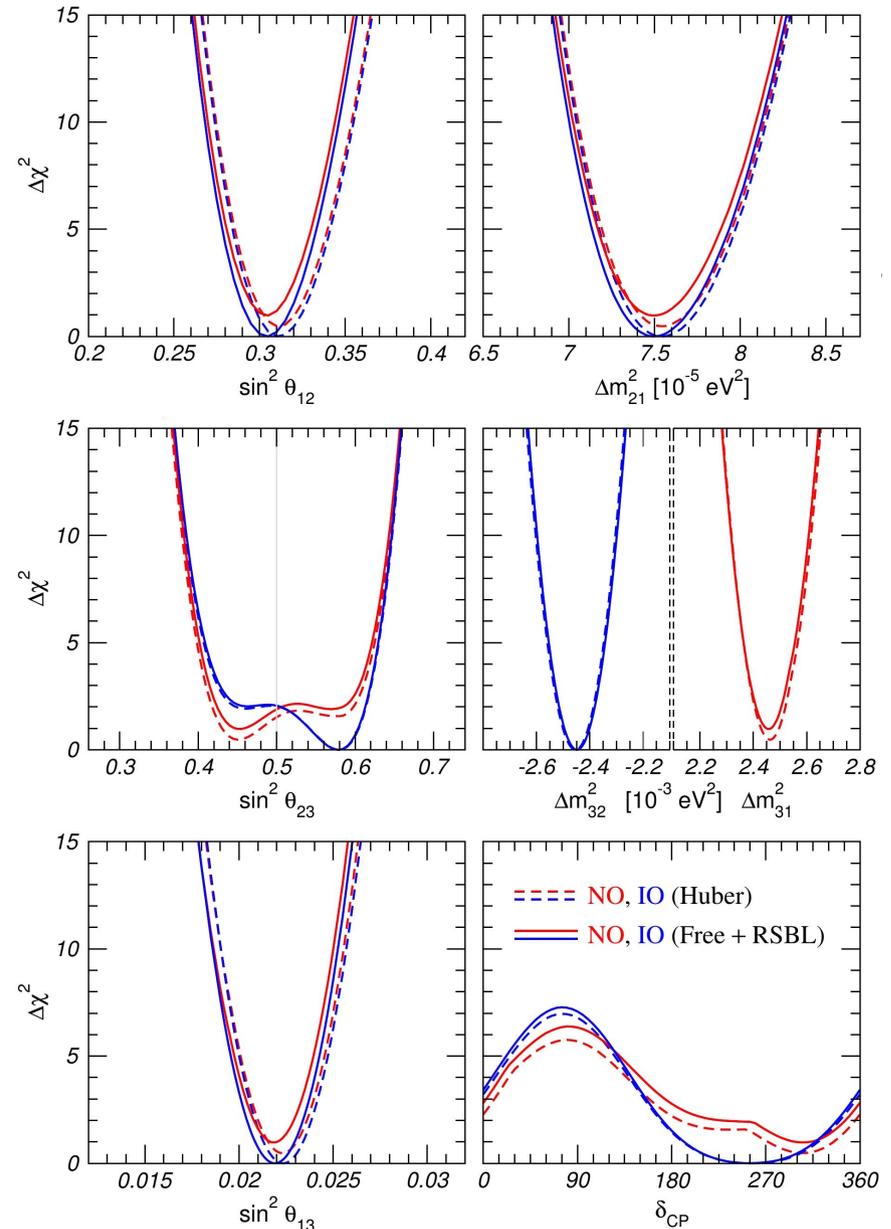
	MI prot/pulse	Energy	Cycle time	Beam power (MW)	Uptime&efficiency		pot/year
	7.50E+13	120	1.2	1.20E+00	0.56		1.10E+21
	7.50E+13	80	0.9	1.07E+00	0.56		1.47E+21
	7.50E+13	60	0.7	1.03E+00	0.56		1.89E+21

- The units: apply (@norm) factor of  $1.017718 \times 10^{17}$  to get the right units
- FD Mass: 40 kt
- Exposure: 6yrs ( $3\nu/3\bar{\nu}$ ;  $\sim 300$  ktMWyr) or 12yrs ( $6\nu/6\bar{\nu}$ ;  $\sim 600$  ktMWyr)
- Energy ranges:  $0.5 < E_{\nu}^{\text{reco}} < 20$  GeV;  $0 < E_{\nu}^{\text{true}} < 120$  GeV

# Study Parameters: Oscillations

- Nominal oscillation parameter values:
  - NuFit 2014 numbers  
<http://www.nu-fit.org/?q=node/92>
  - $\delta_{cp} = 0.0$
  - $\theta_{12} = 0.5843$
  - $\theta_{23} = 0.738$  (0.864 IH)
  - $\theta_{13} = 0.148$
  - $\Delta m_{21} = 7.5 \times 10^{-5}$
  - $\Delta m_{21} = 2.457 \times 10^{-3}$   
 (-2.449  $\times 10^{-3}$  IH)

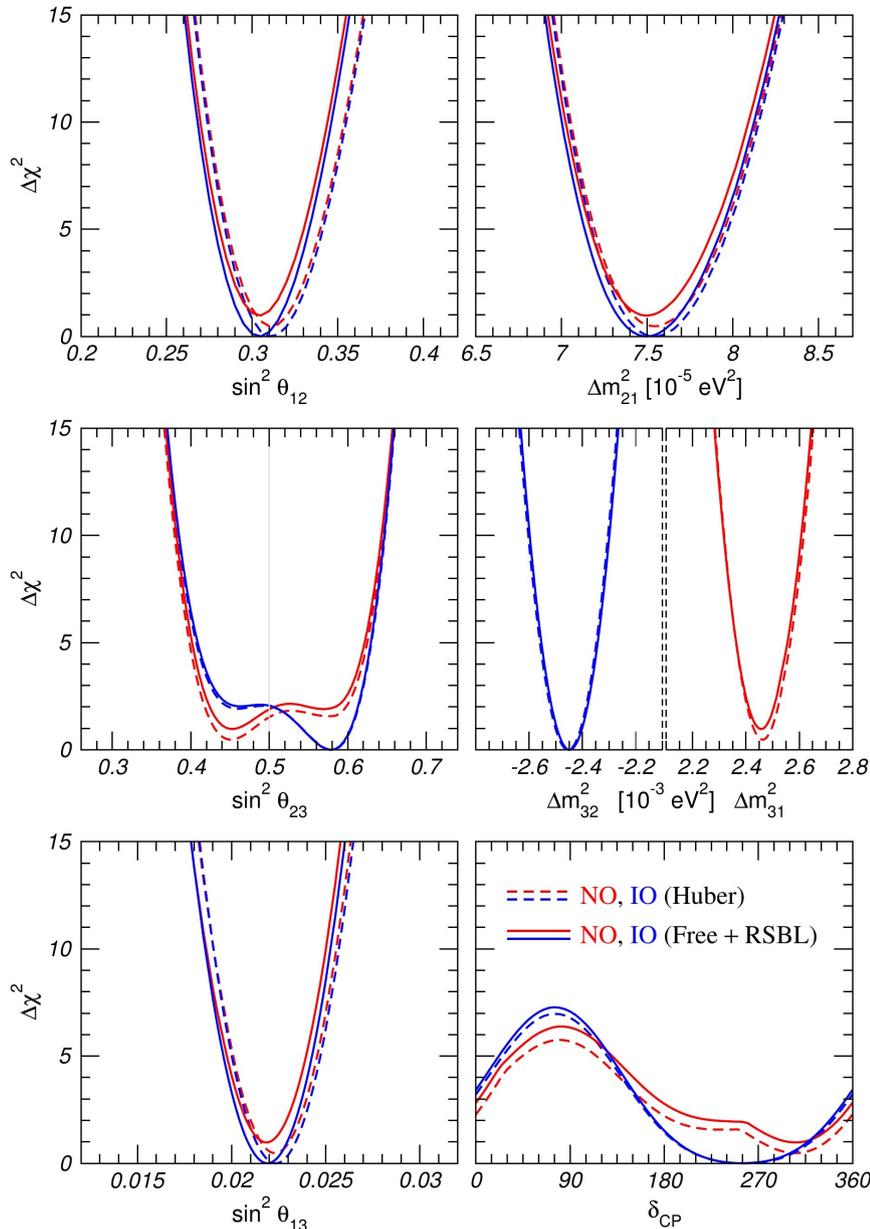
NuFIT 2.0 (2014)



Note: Same inputs as the LOI

# Study Parameters: Oscillations

NuFIT 2.0 (2014)



- Oscillation parameter uncertainties:
    - Assume gaussian uncertainties
    - Define  $1\sigma$  range as  $1/3^{\text{rd}}$  of the  $3\sigma$  range
    - No prior on  $\delta_{\text{CP}}$
    - $\delta\theta_{12} = 2.3\%$
    - $\delta\theta_{23} = 5.9\%$  (4.9% IH)
    - $\delta\theta_{13} = 2.5\%$
    - $\delta\Delta m_{21} = 2.4\%$
    - $\delta\Delta m_{21} = 2.0\%$  (1.9% IH)
- Highly Asymmetric

# Study Parameters: Non-Oscillation Systematics

- FHC/RHC  $\nu_\mu/\bar{\nu}_\mu$  Disappearance
    - Signal  $\nu_\mu + \bar{\nu}_\mu$ : 5%
    - Background NC: 10%
    - Background  $\nu_\tau/\bar{\nu}_\tau$ : 20%
  - FHC/RHC  $\nu_e/\bar{\nu}_e$  Appearance
    - Signal  $\nu_e + \bar{\nu}_e$ : 2%
    - Background  $\nu_e + \bar{\nu}_e$ : 5%
    - Background NC +  $\nu_\mu + \bar{\nu}_\mu$ : 5%
    - Background  $\nu_\tau/\bar{\nu}_\tau$ : 20%
  - Consider FHC and RHC beams separately
- \* See the LBNE SciOp for more details
- Justifications
    - Flux known to  $\sim 5\%$ ;  $\nu_\mu/\bar{\nu}_\mu$  constrains  $\nu_e/\bar{\nu}_e$
    - Assume FHC and RHC are uncorrelated
    - Additional 2% on  $\nu_e$  signal for  $E(\sigma(\nu_e)/\sigma(\nu_\mu))$
    - Large uncertainty on  $\sigma(\nu_\tau)/\sigma(\nu_\mu)$
    - Uncertainty on  $\pi^\pm/\pi^0$  production for NC-like backgrounds
  - Correlations
    - Signal norm. params. for each sample **Uncorrelated** (4 total)
    - NC correlated for FHC+RHC  $\nu_\mu/\bar{\nu}_\mu$  dis.
    - NC +  $\nu_\mu + \bar{\nu}_\mu$  corr. for FHC+RHC  $\nu_e/\bar{\nu}_e$  app.
    - $\nu_\tau$  production correlated across all samples
    - Beam in FHC and RHC  $\nu_e$  uncorrelated 8

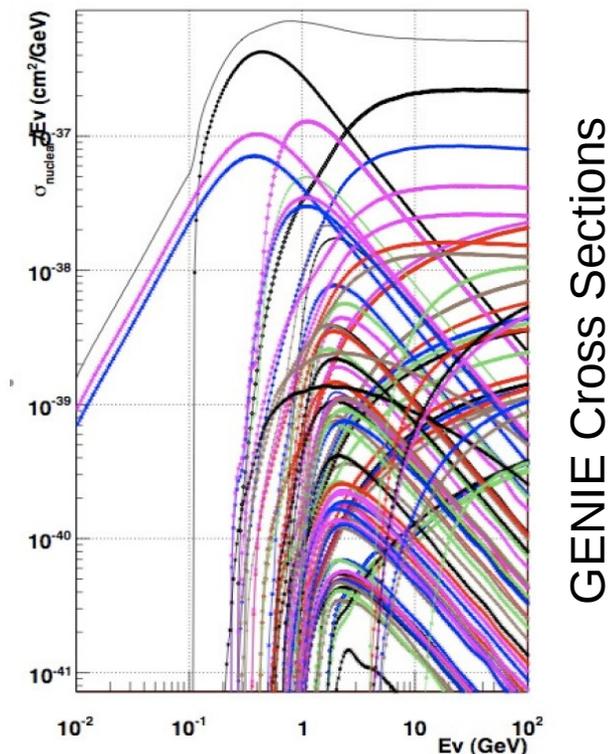
# Study Parameters: Variations

- Beams
  - Variations in beam design
  - See talk from the beam group
- Run Plan
  - Detector construction schedule
  - Beam upgrades
  - $\nu/\bar{\nu}$  run ratio
- Oscillation parameter assumptions
  - Future uncertainties estimates
  - Mass Hierarchy
  - $\theta_{23}$  Octant
- Changes to oscillation parameter central values for:  $\theta_{13}$ ,  $\theta_{23}$ ,  $\Delta m^2_{31}$
- Non-oscillation systematics
  - Altered normalization priors
  - Shape systematics
  - Detailed systematics
    - Fast MC + MGT
    - GENIE cross section reweights
    - Beam optics
  - A lot more to do, but not on CD1r timescales

# GLOBES Inputs: Flux and Cross Sections

- Nominal Flux
  - 80 GeV protons
  - 204m x 4m He filled decay pipe
  - 230kA horn current
  - 1.07 MW
  - $1.47 \times 10^{21}$  POT/yr
  - Be target, -45 cm from horn 1
- Multiple alternate fluxes
  - Will have the 4 flux options implemented
  - Easy to add new ones

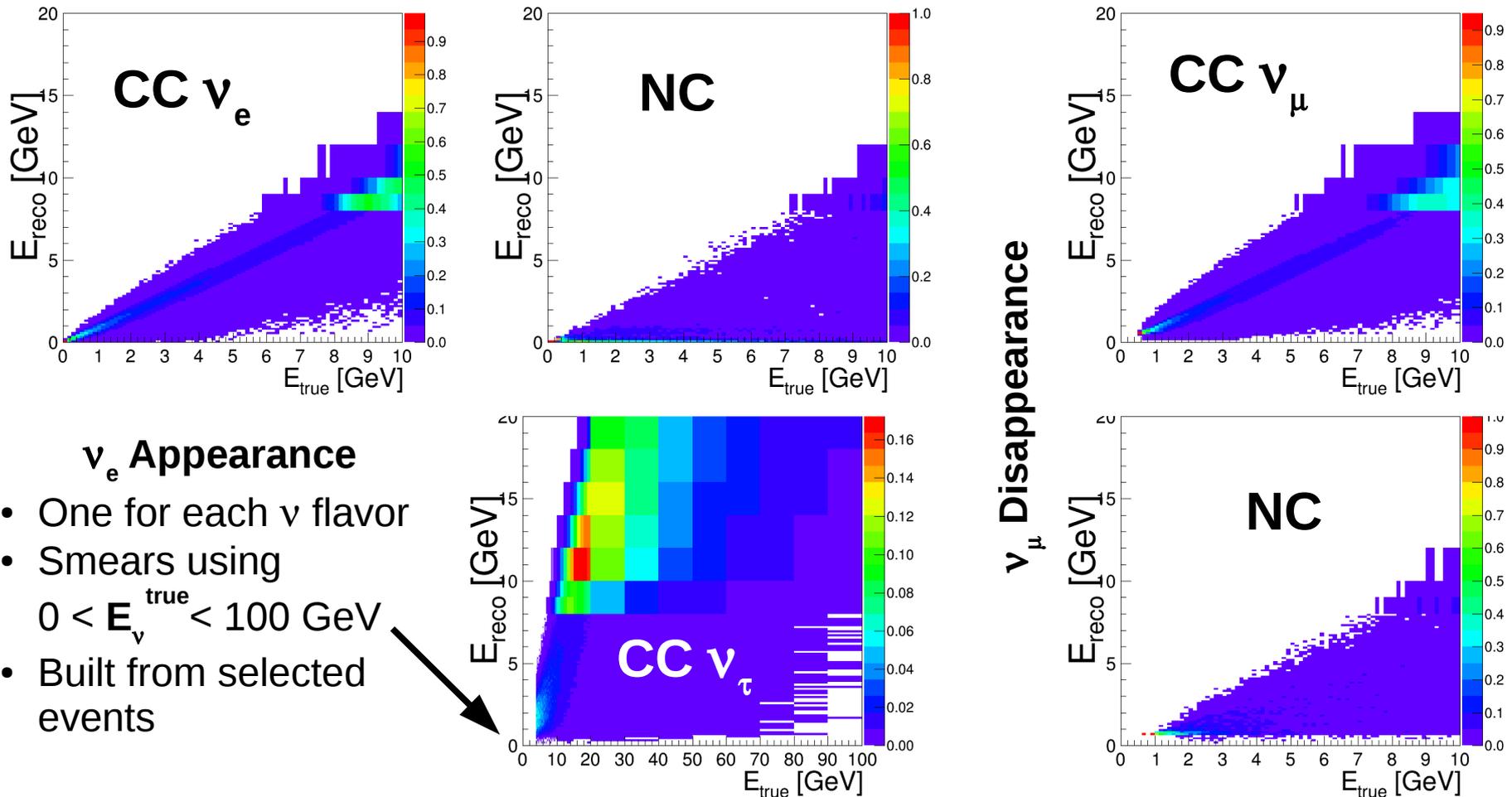
- Cross sections
  - GENIE 2.8.4
  - Inclusive CC and NC
  - $\nu_e$ ,  $\bar{\nu}_e$ ,  $\nu_\mu$ ,  $\bar{\nu}_\mu$ ,  $\nu_\tau$ , and  $\bar{\nu}_\tau$



# GLOBES Inputs:

## Energy Smearing ( $E_\nu^{\text{true}} \rightarrow E_\nu^{\text{reco}}$ )

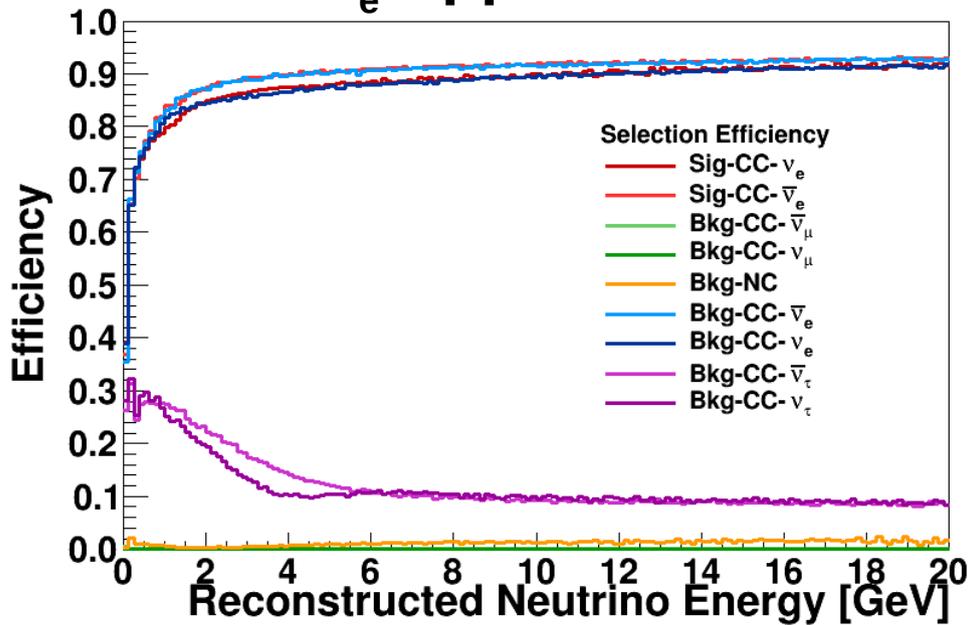
- Generated using the Fast MC
- Plots  $E_\nu^{\text{reco}}$  vs  $E_\nu^{\text{true}}$  and normalize bins of  $E_\nu^{\text{true}}$



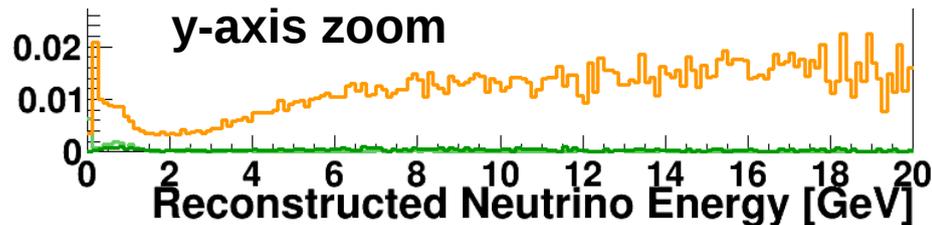
# GLOBES Inputs: Efficiencies

- Generated using the Fast MC
- Ratio: (Fact MC selected events) / (total events)

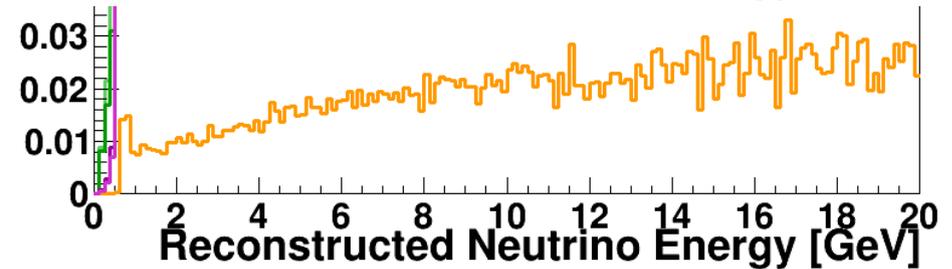
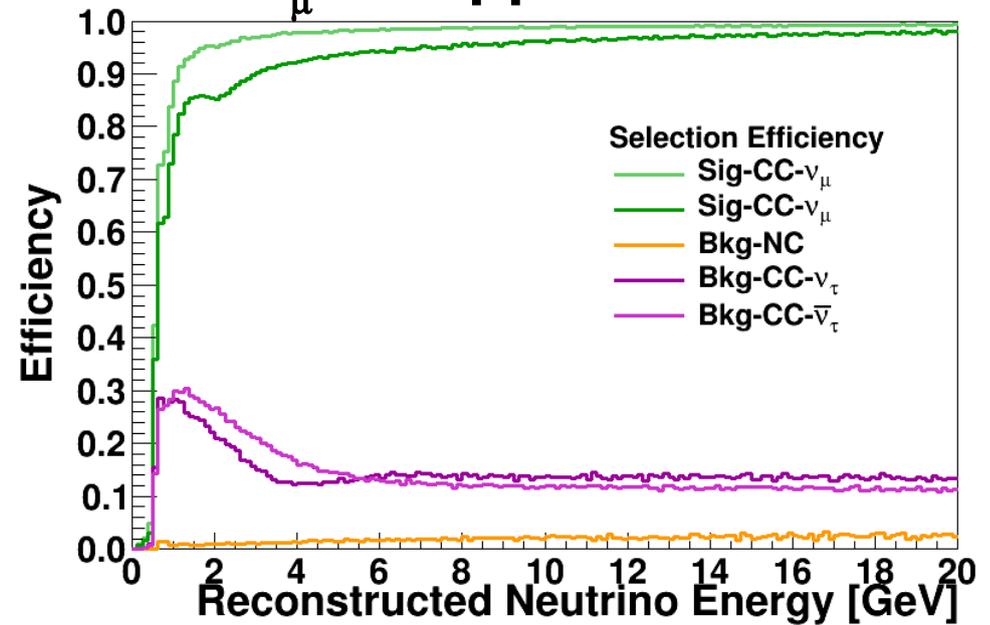
## $\nu_e$ Appearance



### y-axis zoom



## $\nu_\mu$ Disappearance



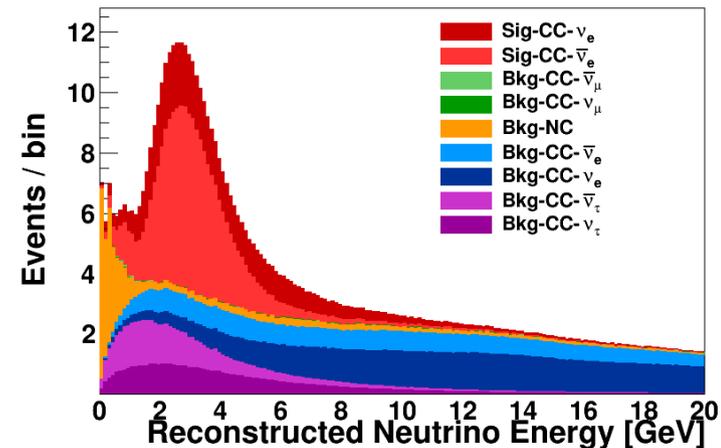
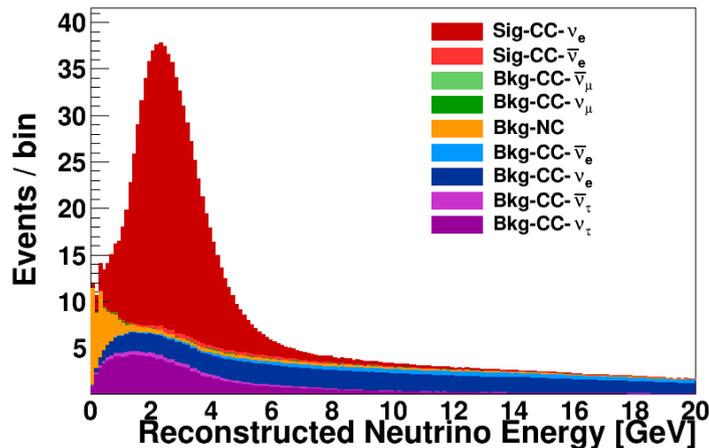
# GLOBES Inputs: Efficiencies – We Can Do Better!

Use transverse momentum cuts (kNN based discriminant)

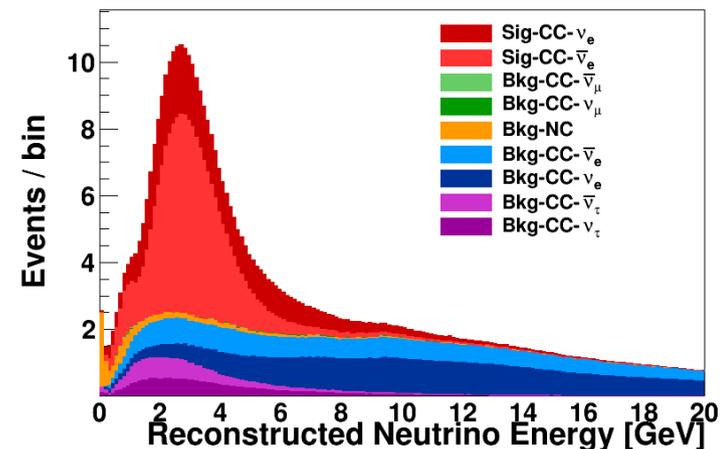
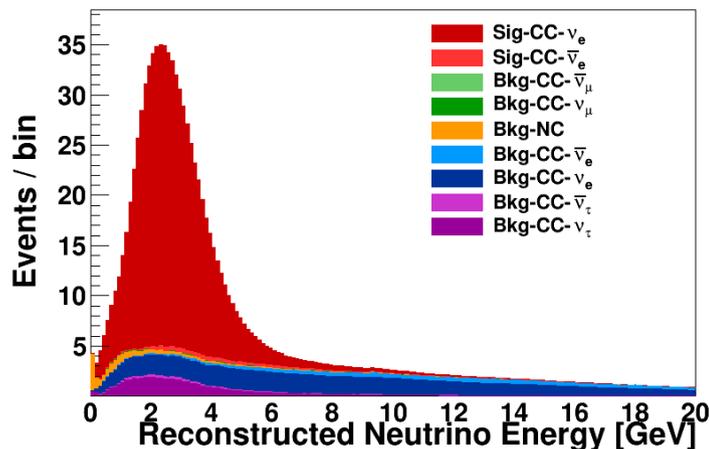
$\nu_e$  Appearance

$\bar{\nu}_e$  Appearance

No Cuts



Both Cuts  
(NC and  $\nu_\tau$ )

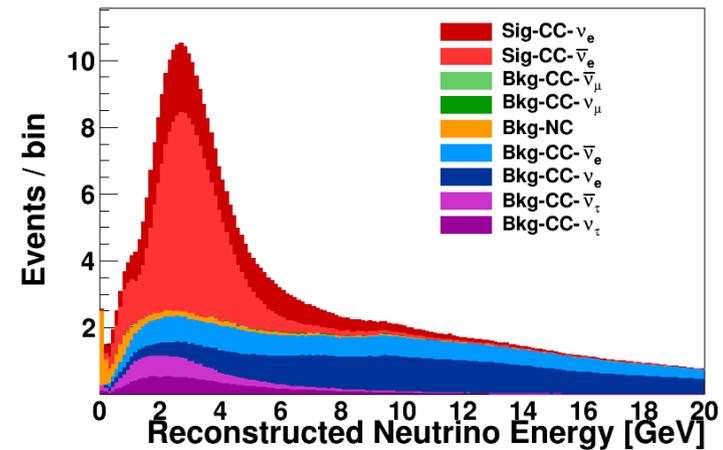
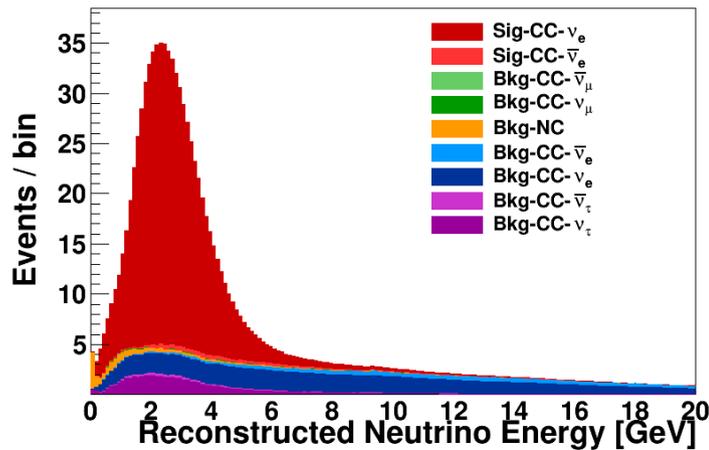


# Energy Spectra Comparisons

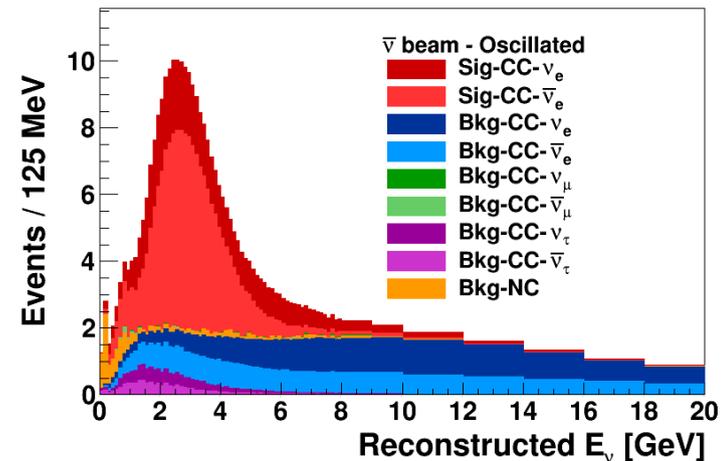
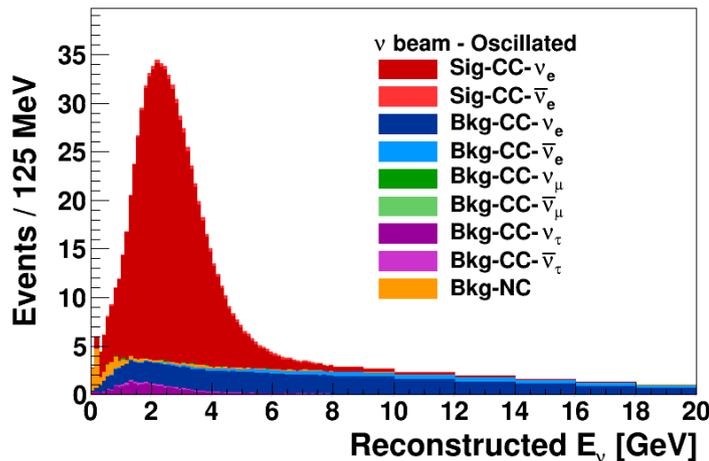
$\nu_e$  Appearance

$\bar{\nu}_e$  Appearance

Fast MC



GLOBES



**Notes:** Different fluxes, and different oscillation parameters

# How to Use this GLoBES Setup

- There is a tarball of all the required files on docdb
  - Must be used from the /lbne/app/ area at FNAL
  - GLoBES config uses relative paths for inputs
  - Some included (.inc) files have absolute paths set for certain inputs
- Copy default GLoBES configuration and change:
  - Fluxes
  - Efficiencies (w/ and w/o transverse momentum cuts)
  - Oscillation parameter central values and uncertainties
  - Non-oscillation systematic uncertainties
  - Run plan (run time,  $\nu/\bar{\nu}$  ratio, detector mass)

# Thinking Forward

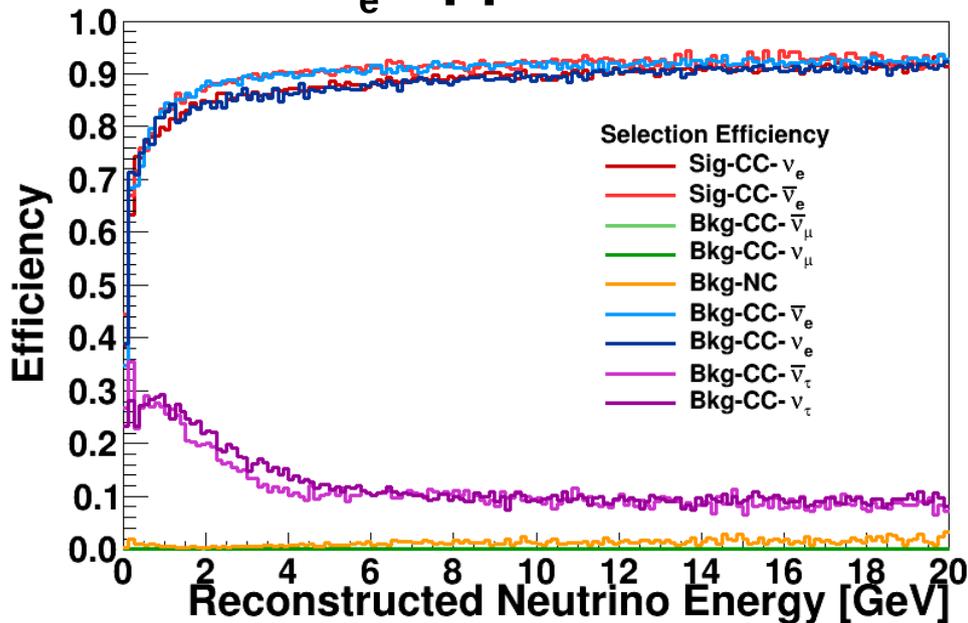
- Since we have detailed inputs, can we move beyond GLoBES?
  - Yes, but it will take some work
  - More study variations
  - Inclusion of detailed non-oscillation systematics
  - Use of oscillation parameter  $\Delta\chi^2$  priors
- See talk in the Systematics Session for more details

# Backup Slides

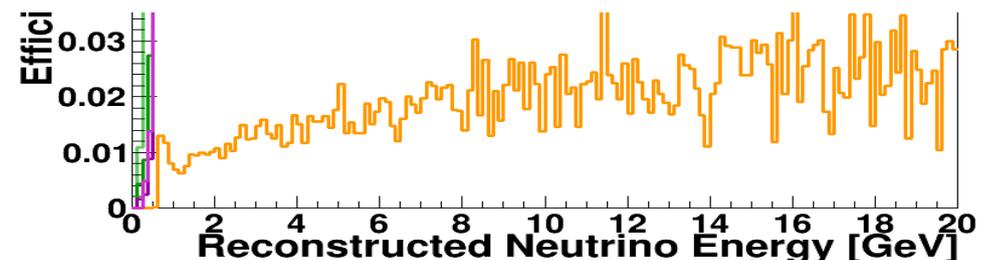
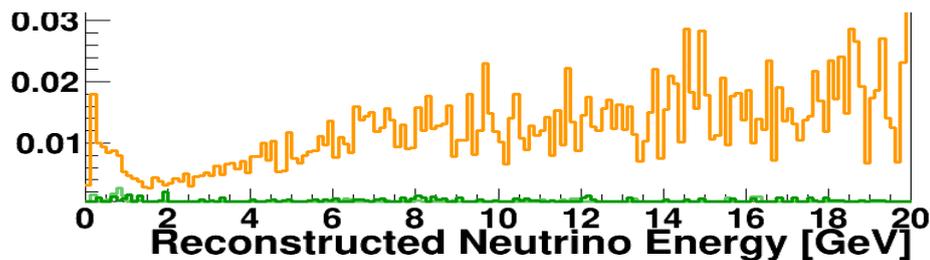
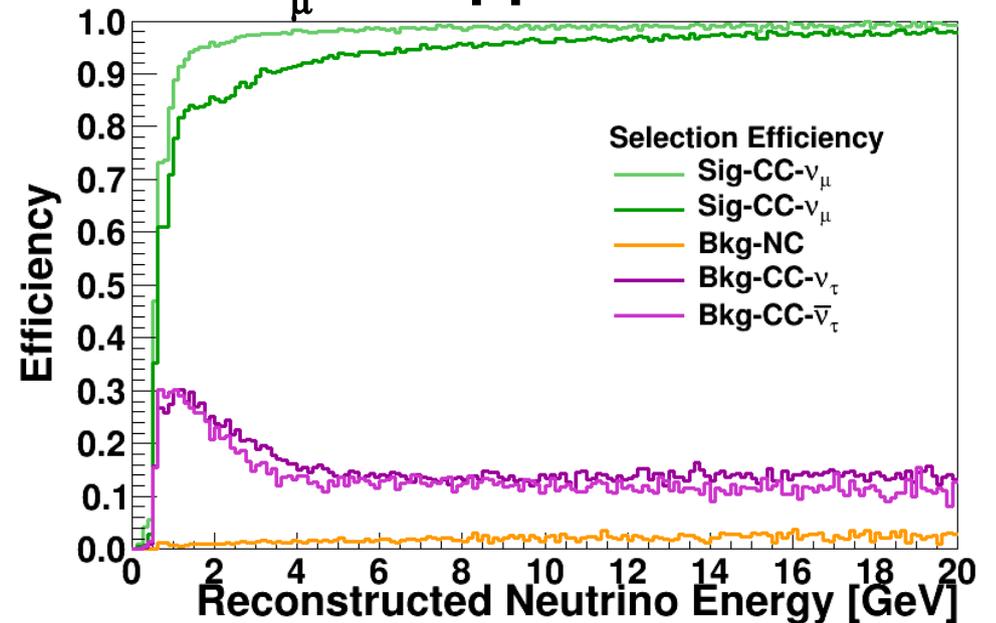
# GLOBES Inputs: Efficiencies

- Generated using the Fast MC
- Ratio: (Fact MC selected events) / (total events)

$\bar{\nu}_e$  Appearance



$\bar{\nu}_\mu$  Disappearance

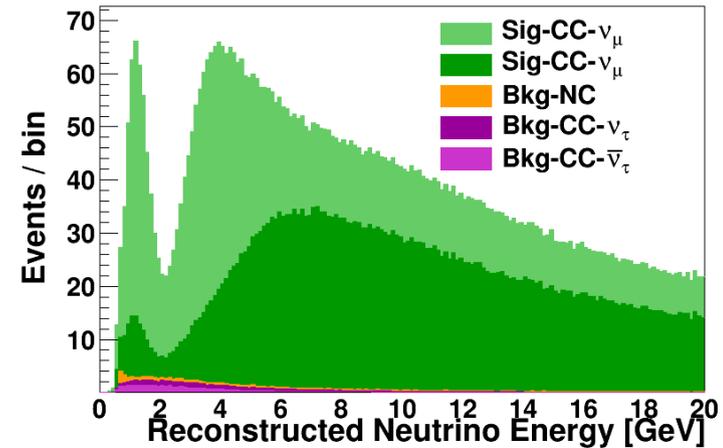
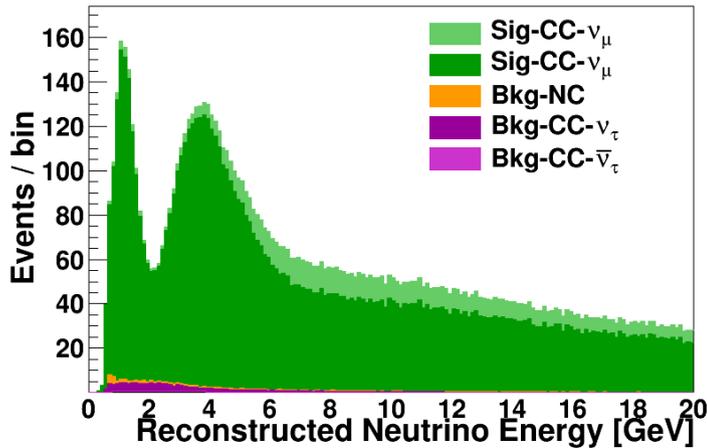


# Energy Spectra Comparisons

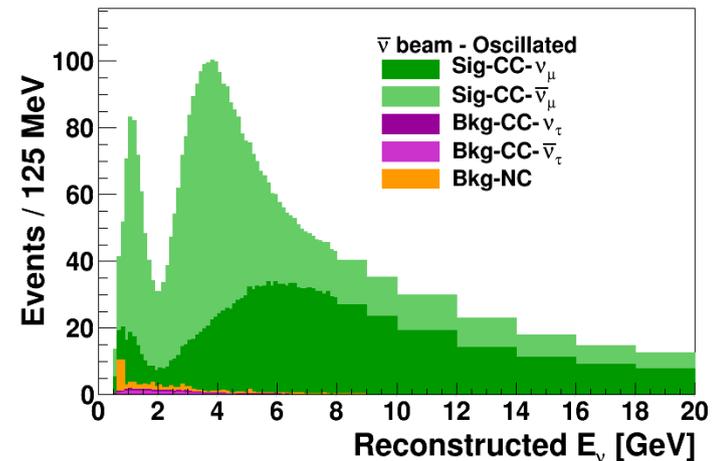
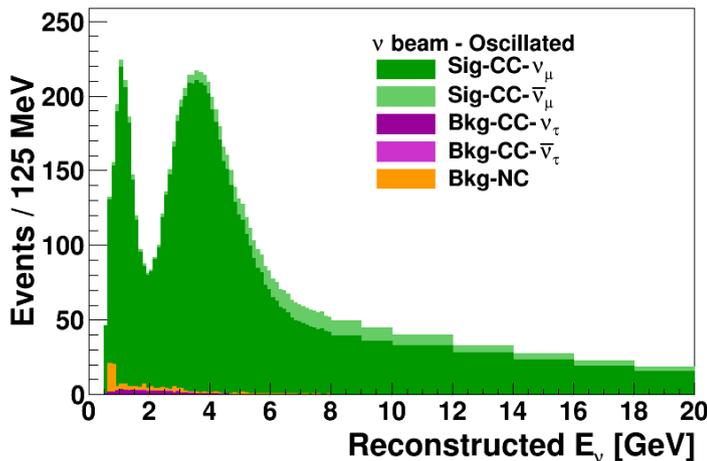
$\nu_\mu$  Disappearance

$\bar{\nu}_\mu$  Disappearance

Fast MC



GLOBES

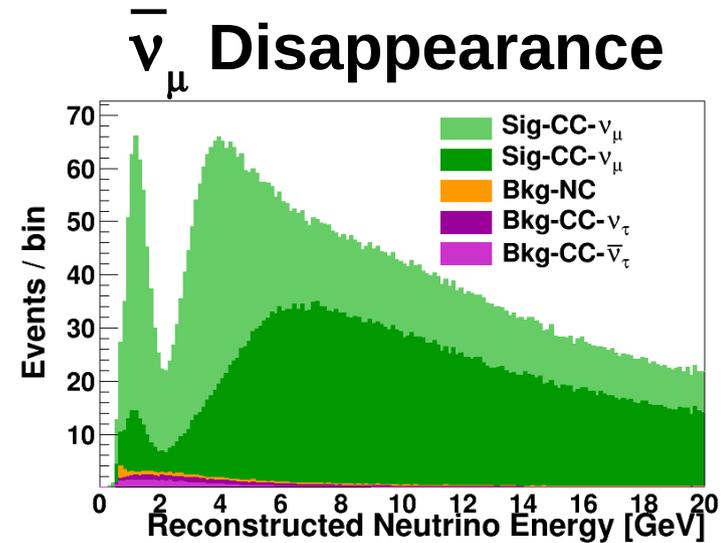
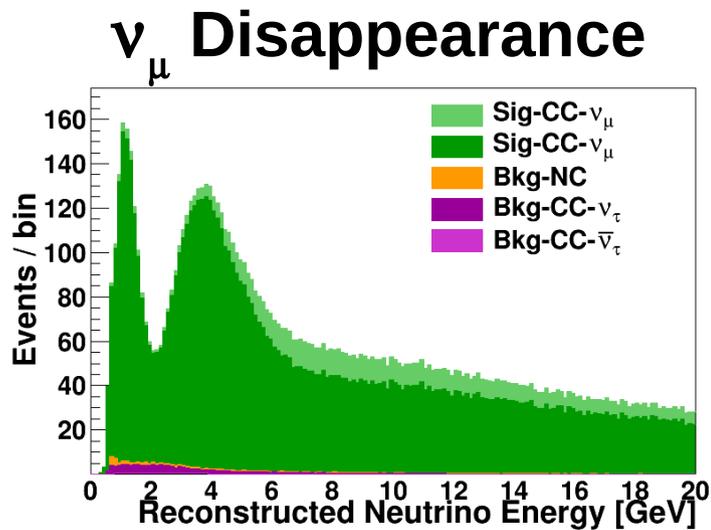


**Notes:** Different fluxes, and different oscillation parameters

# GLOBES Inputs: Efficiencies – We Can Do Better!

Use transverse momentum cuts (kNN based discriminant)

No Cuts



Both Cuts  
(NC and  $\nu_\tau$ )

